Homework 24

## Centripetal acceleration.

We already know that in uniform circular motion, the acceleration vector is inward. How to calculate the magnitude of the centripetal acceleration?


Imagine that an object is moving along a circle from point $A$ to point $B$ (see the picture above). The speed of the object is the same in both points (since it is uniform circular motion). The velocity vectors are different. As the object travels from point 1 to point 2 , its velocity vector turns. We know that at the circular motion the velocity at any point is directed along the tangent line to this point. The tangent line is perpendicular to the radius drawn from the center to the "tangent" point. Simply speaking, each black arrow in the figure above is perpendicular to the corresponding green arrow. So, the angle $\Delta \alpha$ between the velocities $\boldsymbol{V}_{1}$ and $\boldsymbol{V}_{2}$ is equal to the angle $\Delta \boldsymbol{\alpha}$, swept by the moving object.

As we remember, acceleration is the change in velocity divided by the time, required for this change. Change in velocity $\boldsymbol{\Delta V}$ is shown by the red arrow in the picture above. To find it we have to subtract vector $\boldsymbol{V} \mathbf{1}$ from the vector $\boldsymbol{V}$. To do that we will prepare the vector $\mathbf{- V} \boldsymbol{1}$ which has the same length as $\boldsymbol{V} \mathbf{1}$, but its direction is opposite. Then we add $-\boldsymbol{V 1}$ to $\boldsymbol{V} 2$.

We have to find the length of the red arrow $\Delta \mathrm{V}$ and divide it by the time $\Delta t$ which is required for the object to travel from point $A$ to point $B$.

$$
\begin{equation*}
a=\frac{\Delta V}{\Delta t} \tag{1}
\end{equation*}
$$

To calculate $\Delta \mathrm{V}$ we assume that the arc AB is really really small, so the arc AB is very close to a straight line. We can see that

$$
\begin{equation*}
|A B| \approx R \cdot \Delta \alpha \tag{2}
\end{equation*}
$$

It follows from our way to measure the angle. It turns that the formula above is good for any "narrow" isosceles triangle! To find a "short" side we have to multiply one of the "long" sides to the small angle between them. The "narrower" the triangle, the more exact is the formula (2). Let us apply this formula to the triangle formed by $\mathrm{V}_{1}, \mathrm{~V}_{2}$ and $\Delta \mathrm{V}$ :

$$
\begin{equation*}
\Delta V \approx V \cdot \Delta \alpha \tag{3}
\end{equation*}
$$

Let us plug $\Delta \mathrm{V}$ from the formula (3) to the formula (1):

$$
\begin{equation*}
a=\frac{\Delta V}{\Delta t} \approx \frac{V \cdot \Delta \alpha}{\Delta t} \tag{4}
\end{equation*}
$$

But $\Delta \alpha / \Delta t$ is the angular velocity $\omega$. So we can write:

$$
\begin{equation*}
a=V \cdot \omega \tag{5}
\end{equation*}
$$

If we remember that $\omega=\boldsymbol{V} / \boldsymbol{R}$, the magnitude of the acceleration can be written as

$$
a=\frac{V^{2}}{R}
$$

where $R$ is the radius of the circle. So we can see that the faster we move and the sharper we turn the more acceleration we experience and the more centripetal force we need to complete the turn.
Now, as long as we know that an object moves along a circle with a radius R with a constant speed V, we can immediately write its acceleration using the formula (6)!

Problems:

1. Compare the centripetal acceleration of the Earth center due to the orbital motion around the Sun with that of an object at the equator due to Earth rotation. The Earth's orbit radius is $150,000,000 \mathrm{~km}$, the Earth radius is 6371 km (let us assume that it is a perfect circle).
