

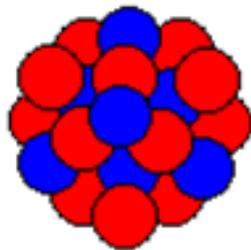
Three Types of Nuclear Reactions

- 1. Radioactive decay** – an unstable nucleus spontaneously emits a small particle of radiation to become a **different isotope** of the same element or a **different element** (such process is called *transmutation*).
- 2. Nuclear Fusion** – the **joining** of two atoms to form a larger one.
- 3. Nuclear Fission** – the **splitting** of an atom into two smaller atoms.

Radioactive Decay

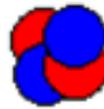
Radioactive decay, also known as radioactivity or nuclear decay, is the process by which a nucleus of an unstable atom loses energy by **emitting ionizing radiation**: ${}^4\text{He}$ (alpha particles), β particles (electrons), γ rays (energetic photons), neutrons.

A heavy nucleus is usually unstable, due to many positive protons pushing apart.



spontaneous decay

alpha particles (He nuclei)



gamma ray



proton



beta particle (electron)



neutron



Radioactive decay is a **random** (*stochastic*) process at the level of single atoms.

Half-Life of Radioactive Isotope

The decay rate of a radioactive isotope is characterized by its **half-life**: the *time it takes for one-half of the atoms of a radioactive material to disintegrate*.

<u>Radioisotope</u>	<u>Half-life</u>
Polonium-215	0.0018 seconds
Bismuth-212	60.5 seconds
Sodium-24	15 hours
Iodine-131	8.07 days
Cobalt-60	5.26 years
Radium-226	1600 years
Uranium-238	4.5 billion years

Naturally Occurring Sources of Radiation

Food:

- **Bananas**, being naturally very high in potassium, consequently have a higher than usual amount of potassium-40, a radioactive isotope.
- The food with the highest concentration of radioactive elements, in this case radium, is the **Brazil nut**.



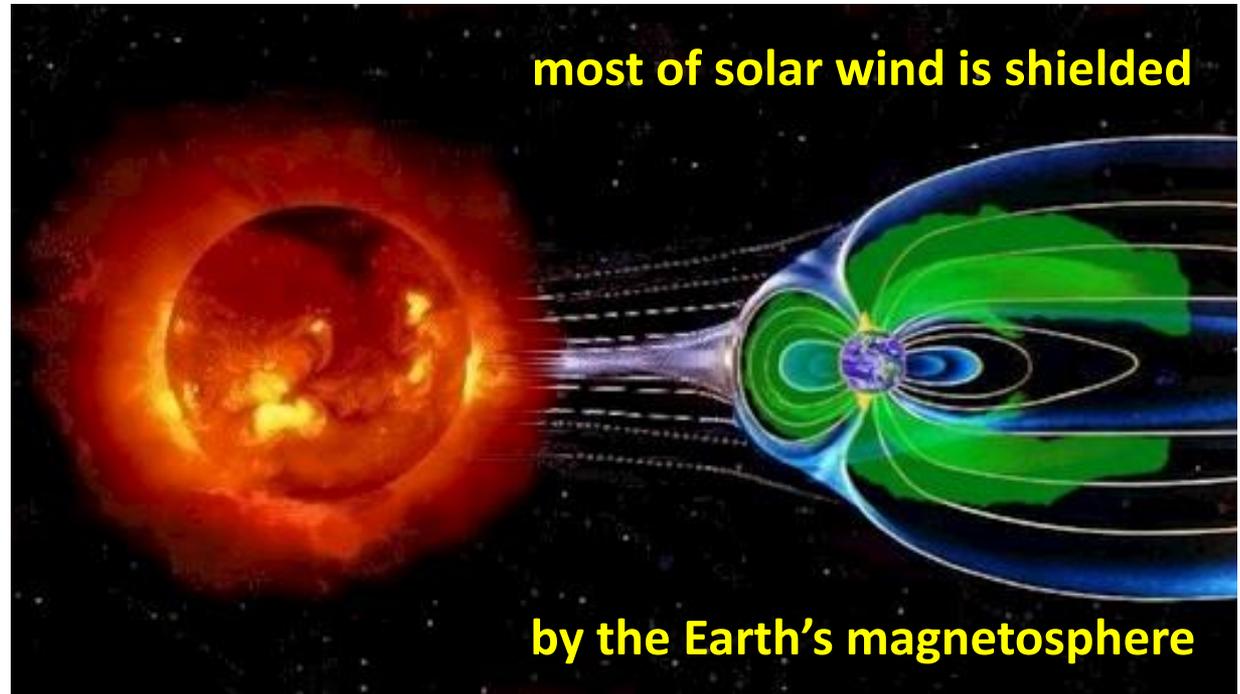
Minerals and materials buried in the earth:

- Most common are potassium-40, uranium-238, and thorium-232 (all with fairly **long half-lives**).
- Additionally, there are small quantities of **shorter-lived** materials (greater activity), such as **radium-226** and **radon-222** (both come as decay products of uranium deposits in the bedrock).
- **Radon**, being a **gas**, can become a problem in some houses and other buildings, seeping in usually through cracks in solid foundations, and accumulating in rooms with poor ventilation.

Naturally Occurring Sources of Radiation

The Sun:

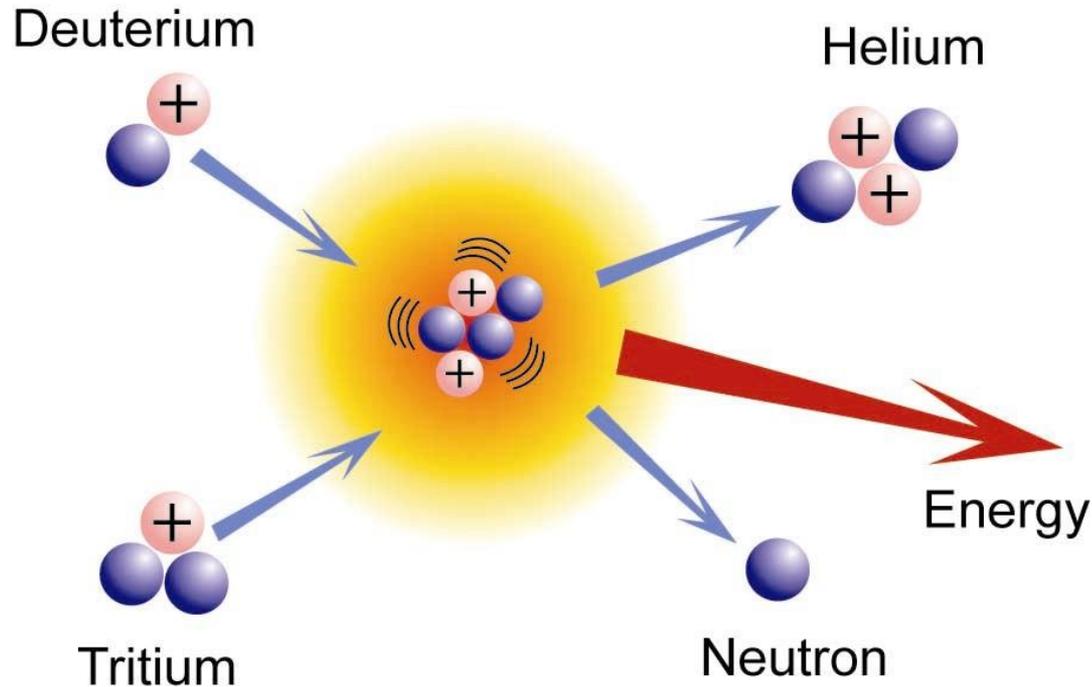
- Powered by a continuous nuclear reaction, main sequence stars give off quite a bit of radiation of every sort!



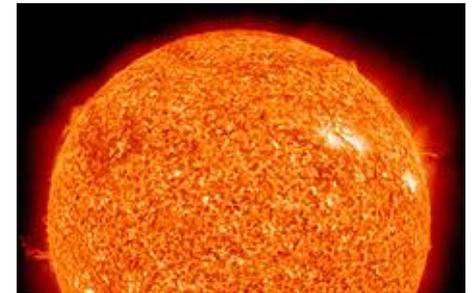
Cosmic radiation:

- Makes up about 14% of the total annual background radiation a person is exposed to over the course of a year.
- The exposure rate is slightly increased by living at higher altitudes, and even more so by air travel (flight crews on long-distance, high-altitude flights tend to accumulate about 30% more annual radiation exposure than the average person!).

Nuclear Fusion



- The fusion of **two nuclei with masses lower than iron generally releases energy**, while the fusion of nuclei **heavier than iron absorbs energy**.



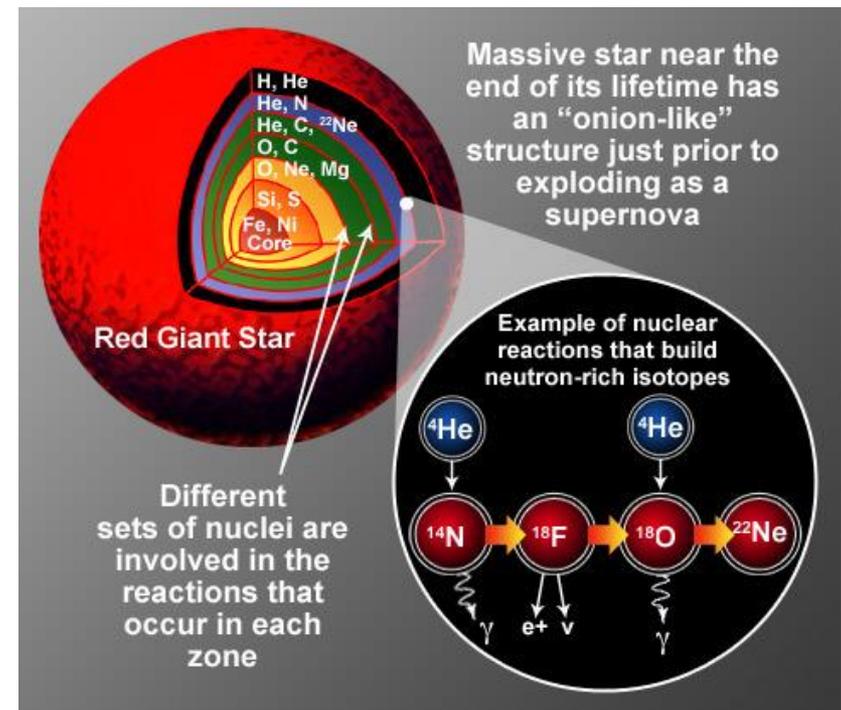
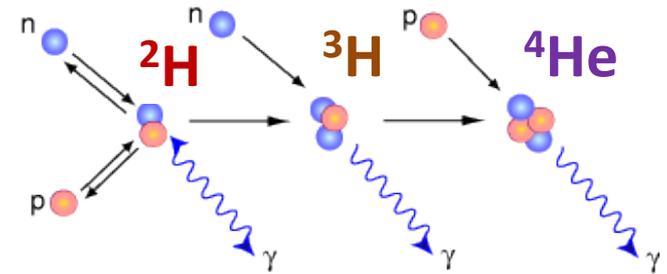
Fusion is the process that powers active stars.

- Fusion reactions have the **greatest energy density**, that is energy per unit of mass, **than any known process** (nuclear fission or chemical reactions).

Nucleosynthesis

Nucleosynthesis is the natural process that **creates new atomic nuclei** from pre-existing nucleons, primarily protons and neutrons:

- Big Bang nucleosynthesis: the first nuclei, **hydrogen and helium**, were formed about *three minutes* after the Big Bang.
- Stellar nucleosynthesis: with the formation of **stars**, heavier nuclei were created from hydrogen and helium, a process that continues today; the **heaviest element** produced by fusion in a normal star is **iron**.
- Supernova nucleosynthesis: production of elements from **iron to uranium** occurs *within seconds* in a supernova explosion.

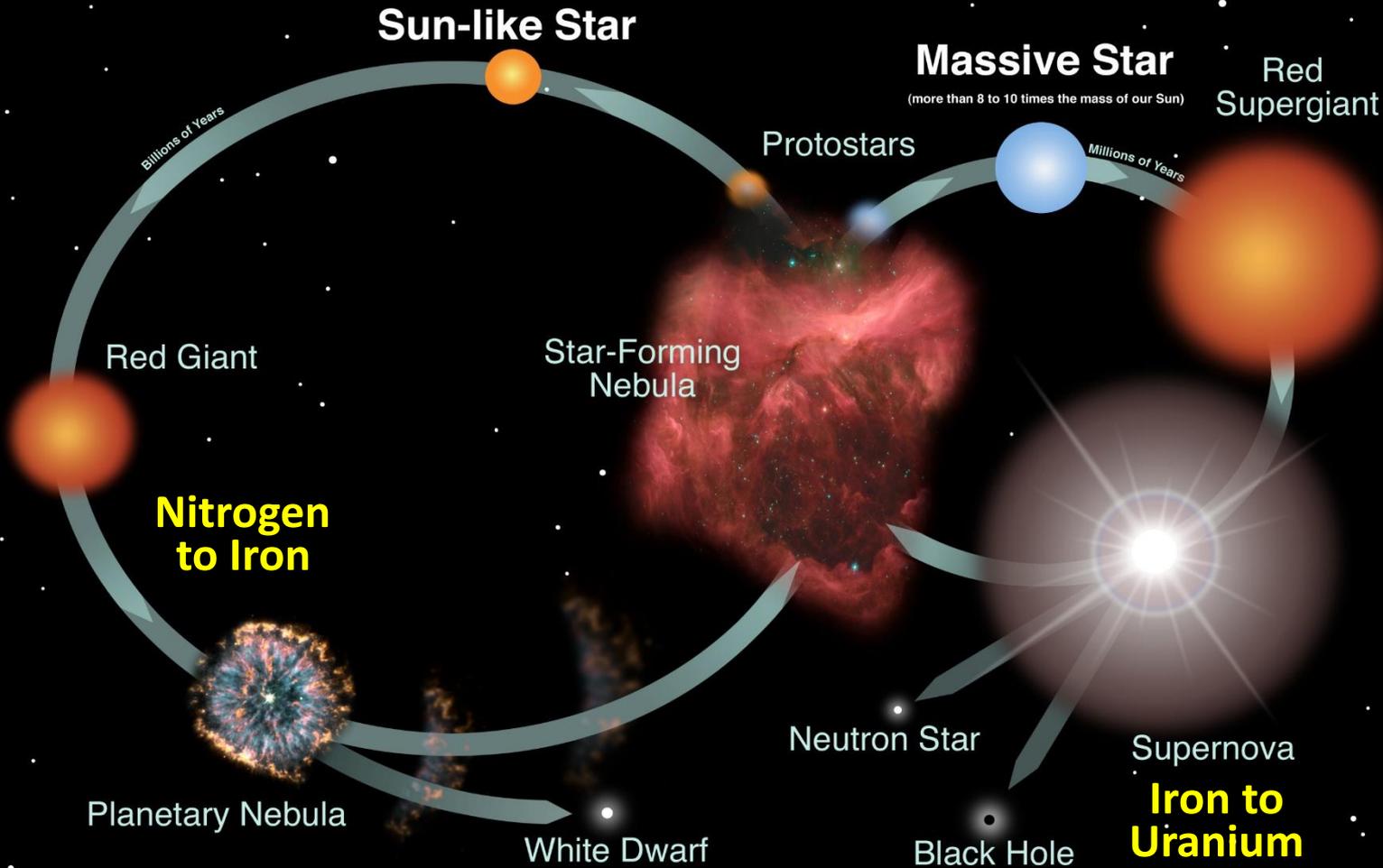


Stellar Recycling



5 minutes after the Big Bang: 75% H and 25% He.

10 billion years of nucleosynthesis: 98% H and He combined, 2% complex elements.



the lives of stars